

Abstract

A high precision local measurement of the Hubble constant H_0 provides direct insight into the physics of the accelerating Universe, a topic that will dominate (astro)physics for decades to come. There is at present an intriguing indication of a significant tension between local measurements of H_0 (SH0ES project: 73.24 ± 1.74 km/s/Mpc, Riess et al 2016, ApJ, 826, 56; see also Freedman et al 2012), which are independent of any cosmological model, and the indirect, highly model-dependent estimate provided by the Planck CMB experiment within the framework of the LambdaCDM concordance cosmological model (67.8 ± 0.9 km/s/Mpc, “Planck 2015 results XIII”, 2016 A&A, 594, 13). If confirmed at a higher significance, this is an unambiguous indication of new physics beyond LambdaCDM, a fact that would have a most profound impact.

In this context, we are carrying out a multi-tiered observational campaign with FLAMES and KMOS at ESO’s VLT to assess at percent level accuracy the impact of the chemical composition of Cepheid stars on the Cepheid-Supernova type Ia distance ladder upon which the SH0ES Riess et al (2016) local measurement of H_0 rests. If not properly accounted for, the effect of chemical composition is expected to bias H_0 at the 1.5-2% level, eating up by itself the entire error budget available for the SH0ES determination of the Hubble constant.

In this talk I will present our project, its status and future perspectives for precision H_0 measurements and their implications.